

◦ Thermalization of a hard parton in a QCD plasma



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Based on:

- » C. Sirimanna et al. [arXiv:2211.15553](https://arxiv.org/abs/2211.15553)
- » Y. Mehtar-Tani, S. Schlichting, and IS [arXiv:2209.10569](https://arxiv.org/abs/2209.10569)
- » S. Schlichting, IS [JHEP 07 \(2021\), 077](https://arxiv.org/abs/2209.10569)



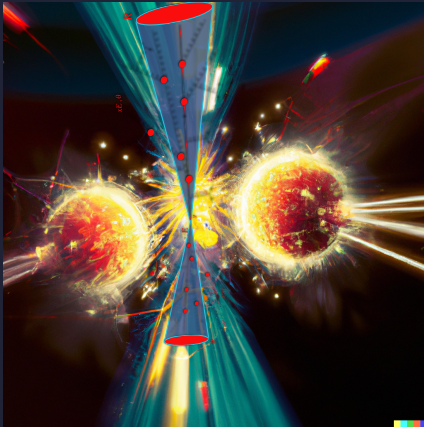
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◉ Outline

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◉ Introduction I

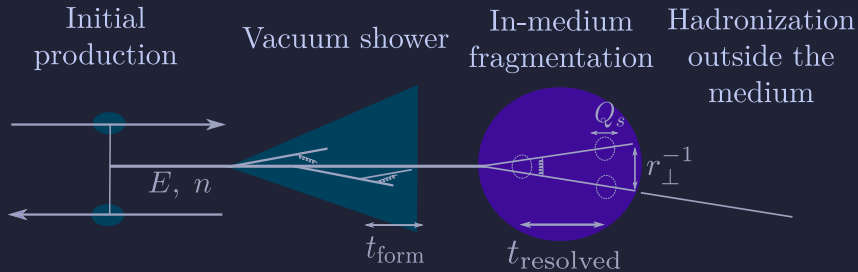


During heavy ion collisions:

- » Deconfined Quark-Gluon Plasma (QGP) is created
⇒ Elliptic flow, Jet quenching
- » Hard partons created at early stages
- » They must traverse the QCD plasma, before reaching the detector
- » By interacting with the medium, they probe the non-equilibrium evolution of the QCD fireball

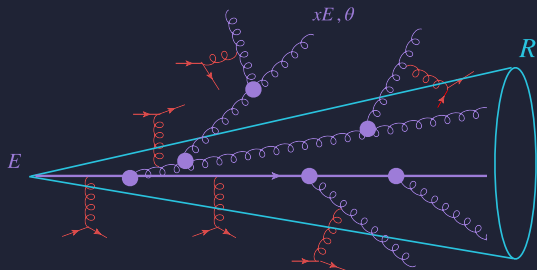
⁰Image generated using DALL-E.

◉ Introduction II



- » The evolution of a hard parton during heavy ion collisions
- » We focus on energy loss and equilibration of hard partons once resolved by the medium

◉ In-Medium Shower



- » Main focus: Hard parton traversing a QCD plasma
- » Understand: Energy cascade, out-of-cone energy loss, medium response and full thermalization of the shower \Rightarrow Important for low energy jets

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◉ Effective Kinetic Description

- » Based on an effective kinetic theory at leading order:

$$p^\mu \partial_\mu f_i(\vec{x}, \vec{p}, t) = C_i[\{f_i\}] , \quad (1)$$

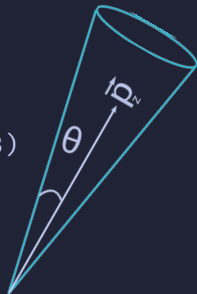
- » The hard partons are described as linearized fluctuations around the equilibrium distribution

$$f(p, t) = n_{\text{eq}}(p, t) + \delta f_{\text{jet}}(p, t) , \quad (2)$$

- » We define energy distribution

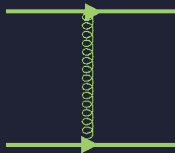
$$D(x, \theta, t) \equiv x \frac{dN_a}{dx d\cos\theta} \sim \frac{\nu_a(N_f)}{E_j} \delta f_{\text{jet}}(p, \theta, t) . \quad (3)$$

- » Exact conservation of energy, momentum and valence charge → Allow us to study the evolution from $\sim E$ to $\sim T$ including full thermalization

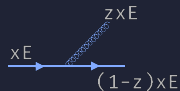


◉ Effective Kinetic Description

» Elastic Scatterings using HTL screened $2 \leftrightarrow 2$



» Collinear Radiation effective $1 \leftrightarrow 2$



$$C[\{f_i\}] = C^{2 \leftrightarrow 2}[\{f_i\}] + C^{1 \leftrightarrow 2}[\{f_i\}]. \quad (4)$$

» Similar to thermalization of QGP in pre-Hydro studies

◉ Elastic Scattering

Collision integral for elastic scattering using HTL screened matrix element

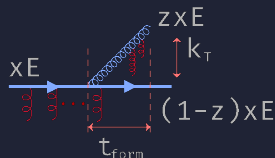
$$C_a^{2\leftrightarrow 2}[\{f_i\}] = \frac{1}{2|\mathbf{p}_1|\nu_a} \sum_{bcd} \int d\Omega^{2\leftrightarrow 2} |\mathcal{M}_{cd}^{ab}(\mathbf{p}_1, \mathbf{p}_2, \mathbf{p}_3, \mathbf{p}_4)|^2 \delta\mathcal{F}(\mathbf{p}, \mathbf{p}_1, \mathbf{p}_2, \mathbf{p}_3, \mathbf{p}_4) \quad (5)$$

Keeping track of the medium response by including the full detailed balance terms

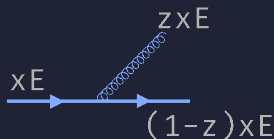
$$\begin{aligned} \delta F(\mathbf{p}_1, \mathbf{p}_2, \mathbf{p}_3, \mathbf{p}_4) = & \delta f_a(\mathbf{p}_1)[\pm_a n_c(\mathbf{p}_3) n_d(\mathbf{p}_4) - n_b(\mathbf{p}_2)(1 \pm n_c(\mathbf{p}_3) \pm n_d(\mathbf{p}_4))] \\ & + \delta f_b(\mathbf{p}_2)[\pm_b n_c(\mathbf{p}_3) n_d(\mathbf{p}_4) - n_a(\mathbf{p}_1)(1 \pm n_c(\mathbf{p}_3) \pm n_d(\mathbf{p}_4))] \\ & - \delta f_c(\mathbf{p}_3)[\pm_c n_a(\mathbf{p}_1) n_b(\mathbf{p}_2) - n_d(\mathbf{p}_4)(1 \pm n_a(\mathbf{p}_1) \pm n_b(\mathbf{p}_2))] \\ & - \delta f_d(\mathbf{p}_4)[\pm_d n_a(\mathbf{p}_1) n_b(\mathbf{p}_2) - n_c(\mathbf{p}_3)(1 \pm n_a(\mathbf{p}_1) \pm n_b(\mathbf{p}_2))] . \end{aligned} \quad (6)$$

◉ Collinear Radiation

- » Multiple scatterings
⇒ induced radiation



- » Resummed into an effective $1 \leftrightarrow 2$



- » Emission controlled by the formation time

$$t_{\text{form}} \sim \frac{z(1-z)xE}{k_T^2} \Rightarrow k_T^2 \sim \hat{q} t_{\text{form}} \quad (7)$$

$$t_{\text{form}} \sim \sqrt{\frac{z(1-z)xE}{\hat{q}}} \Rightarrow \hat{q} \sim \frac{m_D^2}{\lambda_{\text{mfp}}} \quad (8)$$

- » Coherence effects lead to suppression of high energy radiation
⇒ LPM effect

- » $t_{\text{form}} \ll \lambda_{\text{mfp}}$: Medium cannot resolve the quanta until it is formed
- » $t_{\text{form}} \gg \lambda_{\text{mfp}}$: Multiple scatterings act coherently

⁰(Baier, Dokshitzer, Mueller, Peigné, Schiff, Zakharov, Wiedemann, Arnold, Moore, Yaffe..)

◉ Collinear Radiation

Collinear radiation treated using the AMY formalism for an infinite medium (including BH & LPM effects)

$$\begin{aligned} C_g^{g \leftrightarrow gg}[\{D_i\}] = & \int_0^1 dz \frac{d\Gamma_{gg}^g((\frac{xE}{z}), z)}{dz} \left[D_g\left(\frac{x}{z}\right) \left(1 + n_B(xE) + n_B\left(\frac{\bar{z}xE}{z}\right) \right) \right. \\ & + \frac{D_g(x)}{z^3} \left(n_B\left(\frac{xE}{z}\right) - n_B\left(\frac{\bar{z}xE}{z}\right) \right) + \frac{D_g(\frac{\bar{z}xE}{z})}{\bar{z}^3} \left(n_B\left(\frac{xE}{z}\right) - n_B(xE) \right) \Big] \\ & - \frac{1}{2} \int_0^1 dz \frac{d\Gamma_{gg}^g(xE, z)}{dz} \left[D_g(x) (1 + n_B(zxE) + n_B(\bar{z}xE)) \right. \\ & \left. + \frac{D_g(zx)}{z^3} (n_B(xE) - n_B(\bar{z}xE)) + \frac{D_g(\bar{z}x)}{\bar{z}^3} (n_B(xE) - n_B(zxE)) \right], \end{aligned}$$

» Including all the statistical factors which lead to thermalization of soft sector

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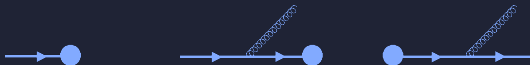
◉ Longitudinal Energy Loss

» Evolution can be divided into three regimes:

1. Initial energy loss: mediated by single gluon radiation
2. Energy cascade: successive emissions lead to an energy cascade
3. Equilibration: Late time decay to soft sector

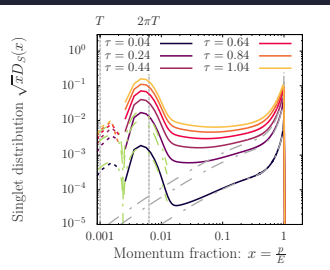
◉ Single Emission

» Initial distribution determined from direct emissions



$$D_q(x, \tau) \simeq \delta(1-x) + \left[x \frac{d\Gamma_{gq}^q(x)}{dz} - \int_0^1 dz \, z \frac{d\Gamma_{gq}^q(x, z)}{dz} \delta(1-x) \right] \tau \quad (9)$$

» Short lived stage \Rightarrow Subsequent emissions lead to energy cascade



◉ Energy Cascade

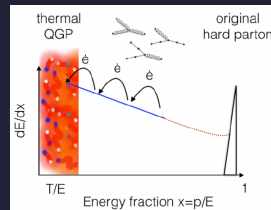
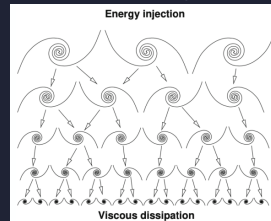
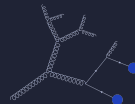
- » Stationary turbulent solution in intermediate energy range $T/E \ll x \ll 1$:

$$D_g(x) = \frac{G}{\sqrt{x}}, \quad D_S(x) = \frac{S}{\sqrt{x}}. \quad (10)$$

- » Fixed point of the differential equation

$$\begin{aligned} C_g[\{D_i\}] &= \int_0^1 dz \frac{d\Gamma_{gg}^g\left(\frac{xE}{z}, z\right)}{dz} D_g\left(\frac{x}{z}\right) - \frac{1}{2} \frac{d\Gamma_{gg}^g(xE, z)}{dz} D_g(x) \\ &\quad + \int_0^1 dz \frac{d\Gamma_{gq}^q\left(\frac{xE}{z}, z\right)}{dz} D_S\left(\frac{x}{z}\right) - N_f \int_0^1 dz \frac{d\Gamma_{qq}^g(xE, z)}{dz} D_g(x), \\ C_S[\{D_i\}] &= \int_0^1 dz \frac{d\Gamma_{gq}^q\left(\frac{xE}{z}, z\right)}{dz} D_S\left(\frac{x}{z}\right) - \frac{d\Gamma_{gq}^q(xE, z)}{dz} D_S(x) \\ &\quad + 2N_f \int_0^1 dz \frac{d\Gamma_{qq}^g\left(\frac{xE}{z}, z\right)}{dz} D_g\left(\frac{x}{z}\right), \end{aligned} \quad (11)$$

$$C_V[\{D_i\}] = \int_0^1 dz \frac{d\Gamma_{gq}^q\left(\frac{xE}{z}, z\right)}{dz} D_V\left(\frac{x}{z}\right) - \frac{d\Gamma_{gq}^q(xE, z)}{dz} D_V(x), \quad (12)$$



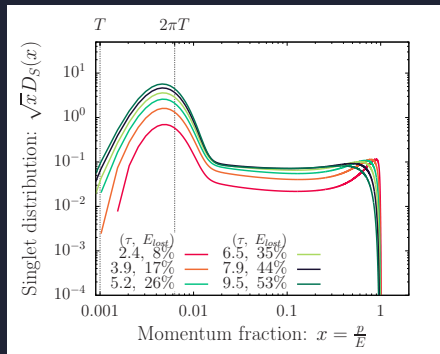
◉ Energy Cascade

- » Stationary turbulent solution in intermediate energy range

$T/E \ll x \ll 1$:

$$D_g(x) = \frac{G}{\sqrt{x}}, \quad D_S(x) = \frac{S}{\sqrt{x}}. \quad (10)$$

- » Existence of a fixed point of the differential equation comes from the fact that the rate behaves as $1/\sqrt{x}E$

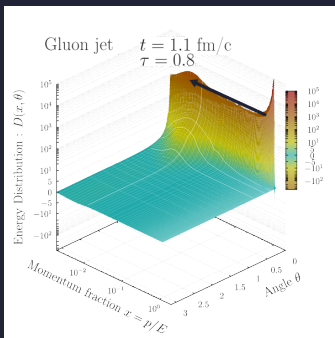


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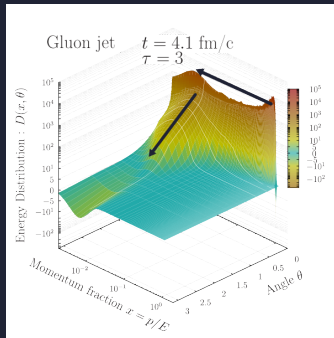
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Thermalization To Large Angles

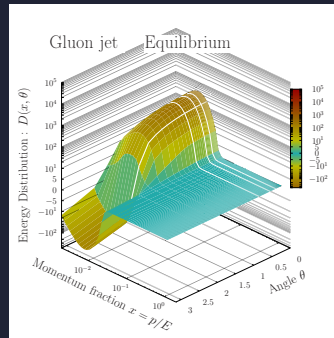
» Energy cascade to soft sector



» Broadening of soft partons to large angle



» Thermalization of soft sector



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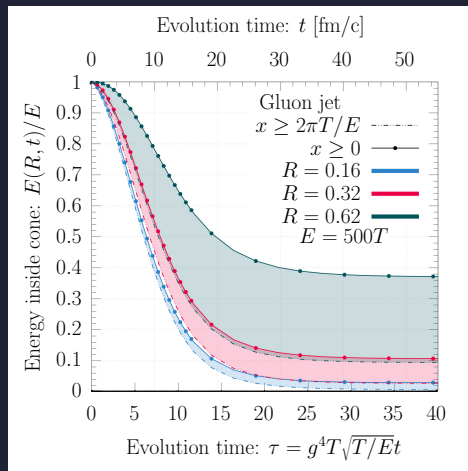
$$^1E/T = 500, \quad g = 2.$$

◉ Cone-Size Dependence

$$\text{---}\bullet\text{---} : E(R, \tau) = \int_0^\infty dx \int_{\cos R}^1 d\cos\theta D(x, \cos\theta, \tau) \quad (11)$$

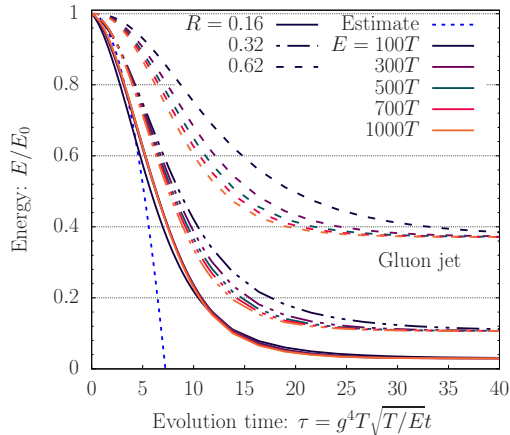
- » Small cone-sizes: Soft sector does not play major role
- » Large cone-sizes: Energy loss display different behavior \Rightarrow Dominated by thermalization

$$\text{---}\bullet\text{---} : E_{2\pi}(R, \tau) = \int_{2\pi T/E}^\infty dx \int_{\cos R}^1 d\cos\theta D(x, \cos\theta, \tau) \quad (12)$$



◉ Sensitivity To Initial Parton's Energy

- » Characteristic time of the turbulent cascade $t_{th} = \frac{1}{\alpha_s} \sqrt{\frac{E}{q}}$
- » Scaling between different initial parton's energy for small cone-sizes
- » Broken for large cone-sizes



◉ Modeling Jet Quenching

» In the presence of QGP, the jet spectrum factorizes

$$\frac{d\sigma}{dp_T} = \int_0^\infty d\epsilon \, P(\epsilon, R) \frac{d\sigma_{\text{vac}}}{dp_T^{\text{in}}} (p_T^{\text{in}} \equiv p_T + \epsilon), \quad (13)$$

» The energy loss probability $P(\epsilon, R)$ is obtained using the BDMPS rate

$$P(\epsilon, R) = \sum_{n=0}^{\infty} \frac{1}{n!} \left[\prod_{i=1}^n \omega_i \frac{dI}{d\omega_i} d\omega_i \right] \delta \left(\epsilon - \sum_{i=1}^n \omega_i \right) \exp \left[- \int_0^\infty d\omega \frac{dI}{d\omega} \right] \quad (14)$$

» Using a Milne transform \Rightarrow Exponential

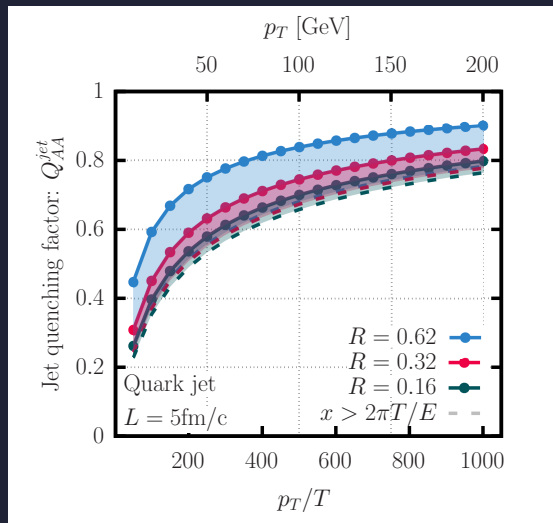
» Taking vacuum spectrum $\frac{d\sigma_{\text{vac}}}{dp_T^{\text{in}}} (p_T^{\text{in}}) \propto p_T^{\text{in}-\alpha}$

$$\frac{d\sigma}{dp_T} = Q(p_T, R) \frac{d\sigma_{\text{vac}}}{dp_T}, \quad \Rightarrow \quad Q(p_T, R) \approx \exp \left[- \int d\omega \frac{dI}{d\omega} \left(1 - e^{-n\omega/p_T} \right) \right] \quad (15)$$

◉ Modeling Jet Quenching

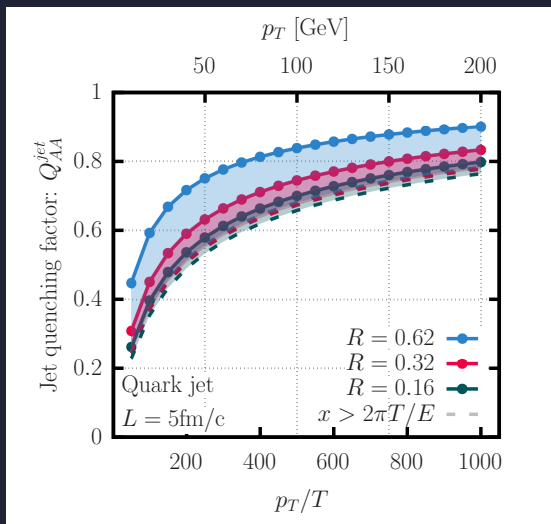
- » The first emission is modeled using BDMPS finite medium rate $\frac{d\Gamma}{d\omega}(P, \omega, t)$ at time t
- » Medium energy loss computed by modeling the energy remaining inside the cone $E(\omega, R, L-t)$ after a time $(L-t)$ in the medium

$$Q(p_T) = \exp \left[\int_0^L dt \int d\omega \times \frac{d\Gamma}{d\omega} \left(1 - e^{-n \frac{\omega}{p_T} \left[1 - \frac{E(\omega, R, \tau = \frac{L-t}{\tau_{th}})}{\omega} \right]} \right) \right]. \quad (16)$$



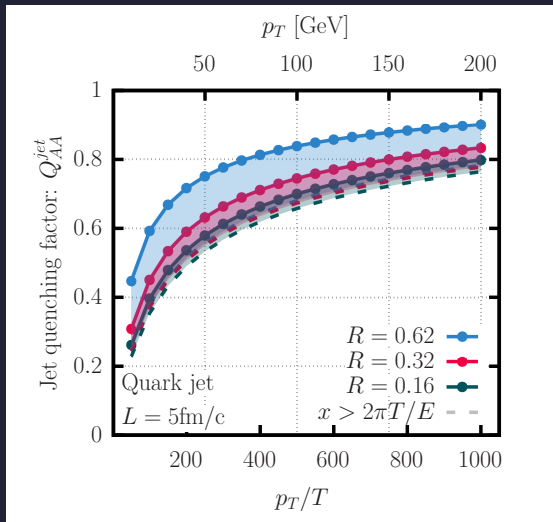
◉ Modeling Jet Quenching

- » Jets with energies $p_T \leq 25\text{GeV}$ lose significant energy in length $L = 5\text{fm}/c$
 \Rightarrow large suppression at low p_T , milder for $p_T \leq 200 - 400\text{GeV}$
- » Negligible contribution of soft fragments to narrow cone
- » Large cone size (≥ 0.3) \Rightarrow Recover energy from the soft sector
 - ▷ Medium response



Modeling Jet Quenching

- ⚠ Over-estimate energy loss since we neglect finite size effects, Work in progress
- ⚠ Requires more refined studies of near-equilibrium physics and jet recoil onto the medium



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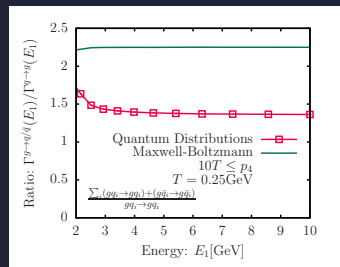
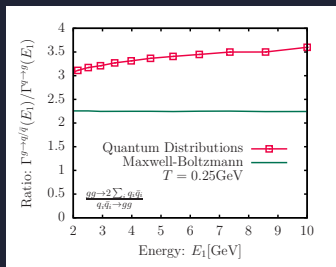
◉ Flavor Conversion

- » Due to scatterings with the medium \Rightarrow Partons change flavor
- » The rate of flavor conversion is determined by the d.o.f. of the medium

◉ Ratio of the rate $\frac{\Gamma_{g \rightarrow q}}{\Gamma_{q \rightarrow g}}$

» Annihilation $g\bar{g} \rightarrow q\bar{q}$

» Scattering $gq_i \rightarrow gq_i$



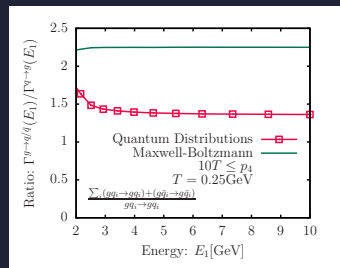
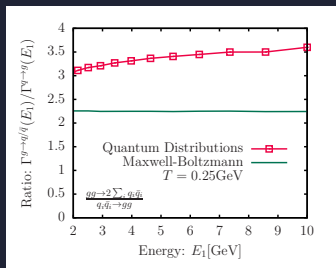
$$\Gamma_{ab \rightarrow cd} = \int \frac{d^3 \mathbf{p}_2}{(2\pi)^3} \frac{d^3 \mathbf{p}_3}{(2\pi)^3} \frac{d^3 \mathbf{p}_4}{(2\pi)^3} f_b(\mathbf{p}_2) [1 \pm f_c(\mathbf{p}_3)] \frac{|\mathcal{M}_{ab \rightarrow cd}|^2}{16 E_1 E_2 E_3 E_4} (2\pi)^4 \delta^{(4)}(p_1 + p_2 - p_3 - p_4),$$

◉ Flavor Conversion

- » The rate of flavor conversion is determined by the d.o.f. of the medium
- » Since gluons have a larger # partons to scatter of (quarks must scatter with other quarks of same flavor) \Rightarrow Gluons are converted to quarks at a higher rate than quarks to gluons

» Annihilation $gg \rightarrow q\bar{q}$

» Scattering $gq_i \rightarrow gq_i$

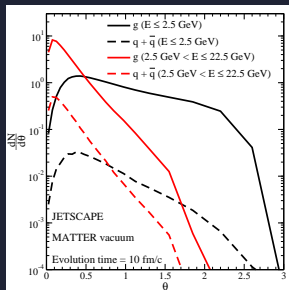


$$\Gamma_{ab \rightarrow cd} = \int \frac{d^3 \mathbf{p}_2}{(2\pi)^3} \frac{d^3 \mathbf{p}_3}{(2\pi)^3} \frac{d^3 \mathbf{p}_4}{(2\pi)^3} f_b(\mathbf{p}_2) [1 \pm f_c(\mathbf{p}_3)] \frac{|\mathcal{M}_{ab \rightarrow cd}|^2}{16 E_1 E_2 E_3 E_4} (2\pi)^4 \delta^{(4)}(p_1 + p_2 - p_3 - p_4),$$

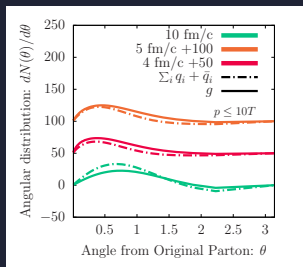
◉ Flavor Composition Of The Soft Sector

◉ $E \leq 2.5 \text{ GeV}$

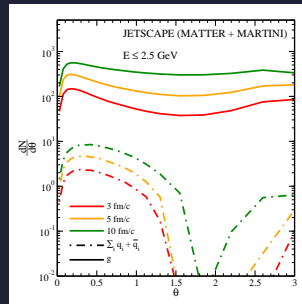
» Vacuum (MATTER)



» Kinetic evolution



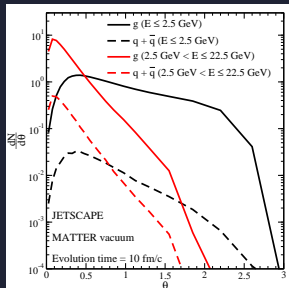
» Vacuum + Energy loss (MATTER + MARTINI)



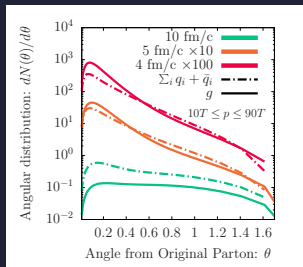
◉ Flavor Composition Of The Shower

$$\textcircled{\circ} \quad 2.5\text{GeV} \leq E \leq 22.5\text{GeV}$$

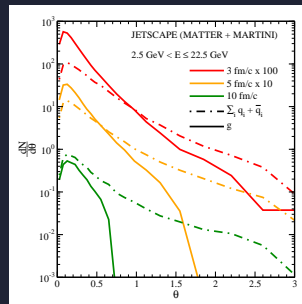
» Vacuum (MATTER)



» Kinetic evolution



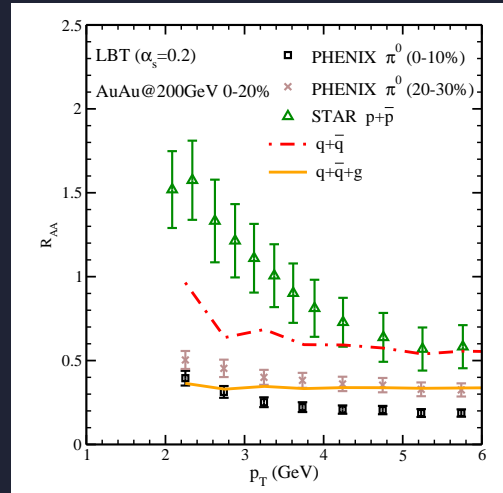
» Vacuum + Energy loss (MATTER + MARTINI)



» Quark content of the in-medium shower is much more quark-like than in vacuum

◉ Flavor Composition Of The Shower

- » An order of magnitude increase in the fermion content of jets due to the medium.
- » New transport coefficients needed to incorporate quark exchange in jet quenching discussion.
- » Increase in fermion content affects conserved charge fluctuations, not energy profile of the jet.
- ⚠ Hadronization introduces own fluctuations and may introduce additional energy loss. (Work in progress)



◉ Conclusion

- » Energy loss is governed by an inverse energy cascade
⇒ driven by successive collinear radiation
- » Energy deposited collinearly at the soft scales rapidly broadens to large angles
- » Kinetic evolution can be seen as a Green's function
⇒ can be convoluted to obtain more evolved evolution
- » Chemistry of jets is highly sensitive to the medium
⇒ several interesting effects at play: baryon enhancement, QGP chemistry, etc.

Thank you for your attention \o/

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